

The role of accelerometer data calibration within the ITSG-Grace2016 release: impact on C20 coefficients

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Outline



- 1) Accelerometer data
- 2) Calibration approach
- 3) Biases & Scale factors
- 4) Impact on C20
- 5) Conclusions & Outlook

Accelerometer data



SuperSTAR accelerometer

- Three-axis electrostatic accelerometer (ONERA)
- Two high-sensitive axes: along-track, radial
- One less-sensitive axis: cross-track

Accelerometer Level-1B data:

- ACC1B data contains instrument bias and scale
- A-priori values from GRACE Technical Note TN-02 (Bettadpur, 2008)
- April 2011: active thermal control was switched off
- Temperature variations correlated with beta prime (β') angle variations
- **Disturbance effects:** thruster firings, heater switches, twangs, magnetic torquer induced accelerations, ... (Flury et al., 2008; Peterseim et al. 2012)





Calibration approach

Accelerometer biases & scale factors:

- Two-step approach:
- a-priori calibration for data screening
- Calibration equation:

 $\mathbf{a}_{\mathrm{cal}} = \mathbf{S} \mathbf{a}_{\mathrm{obs}} + \mathbf{b}$



- Main-diagonal elements
- Shear parameter
- Rotation parameter

(1) Biases:

- Estimation: once per day
- Parameterization: uniform cubic basis splines (UCBS), with a 6h knot interval

(2) Scale factors:

- Estimation: once per day
- Parameterization: fully-populated scale factor matrix
- Off-diagonal elements: non-orthogonality of accelerometer axes (cross-talk), misalignment between SRF and AF

Calibration approach

Accelerometer biases & scale factors:

- Two-step approach:
- a-priori calibration for data screening
- Calibration equation:

 $\mathbf{a}_{cal} = \mathbf{S} \ \mathbf{a}_{obs} + \mathbf{b}$ with $\mathbf{S} = \begin{bmatrix} s_x & \alpha + \zeta & \beta - \epsilon \\ \alpha - \zeta & s_y & \gamma + \delta \\ \beta + \epsilon & \gamma - \delta & s_z \end{bmatrix}$

- Main-diagonal elements
- Shear parameter
- Rotation parameter

1st step calibration – Data pre-processing:

- Calibration parameters: biases and scale factors
- Modeled non-gravitational accelerations are used as reference to estimate calibration parameters
- Enables data screening

2nd step calibration – Gravity field recovery:

- Calibration parameters: biases and scale factors
- Re-estimation of calibration parameters

The same parameterization is used for both steps!





Bias = offset + drift

Bias

- Temperature-induced bias drifts:
 - Related to occasional disabling of heaters (< 2011-04)
 - Related to orbital configuration w.r.t the Sun (> 2011-04)





Bias drifts



- After thermal control switch-off: bias drifts related to orbital configuration
- Heating and cooling of the satellite: cross-track axis shows strongest variations



Bias drifts



- After thermal control switch-off: bias drifts related to orbital configuration
- Heating and cooling of the satellite: cross-track axis shows strongest variations
- Temperature changes highly correlated with beta prime (β') angle variations



Temperature & Beta Prime Angle



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Scale factors



Main diagonal elements:

- Scale factors: along-track (s_x), cross-track (s_y), radial (s_z)
- Non-constant behavior
- High sensitive axes better estimable and less scattered





Scale factors



- Off-diagonal elements xy, xz, yz
 - Shear parameter: α , β , γ
 - Rotational parameter: ζ, ε, δ
- Shear and rotational parameters highly correlated





Temperature-dependency

TU Graz

April 2011 – present:

- Scale factors highly correlated with temperature variations (> 2011-04)
- Temperature variations are absorbed by calibration parameters and map into time-series





Atmospheric density (DTM2013)

- **TU** Graz
- Scale factors better estimable for periods with higher atmospheric densities (i.e. larger nongravitational signal)
- Density variations depend on solar activity, geomagnetic activity and altitude





Beta prime (β ') angle



 Shear parameter: mutual influence among the cross-track and radial axes, due to non-orthogonality of AF and SRF

161-day periodic signal





Altitude



- Interference from other axis components: magnitude dependent on magnitude of the actual non-gravitational accelerations
- Misalignment errors are more significant for lower altitudes where larger atmospheric drag is present



Impact on C20 coefficients

- **TU** Graz
- Fully-populated scale factor matrix: offset w.r.t SLR is reduced (2008-2014)

Impact on C20 coefficients

- **Fully-populated scale factor matrix:** offset w.r.t SLR is reduced (2008-2014)
- Differences increase at the beginning and end of GRACE time-series

c₂₀ + 4.841694552725e-04

Conclusions & Outlook

- **TU** Graz
- GRACE accelerometers are extremely sensitive to temperature variations
- Temperature-induced variations of calibration parameters (biases & scale factors)
- Fully-populated scale factor matrix significantly improves estimates of C20 coefficients
- ACC parameterization also influences:
 - Other low degree coefficients
 - Overall accuracy of monthly gravity field solutions
- Further analysis: ideal parametrization of calibration equation
 - Model not "physically correct"
 - Parameters are likely to absorb other spurious signals

 Article: Klinger, B., Mayer-Gürr, T., 2016. The role of accelerometer data calibration within GRACE gravity field recovery: Results from ITSG-Grace2016. Adv. Space Res. 58, 1597-1609. <u>http://dx.doi.org/10.1016/j.asr.2016.08.007</u>

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Degree amplitudes

- ITSG-Grace2016 (prelim): main-diagonal elements only
- ITSG-Grace2016:

fully-populated scale factor matrix

Modeled non-conservative accelerations

